

# HERO: High-Energy Replicated Optics

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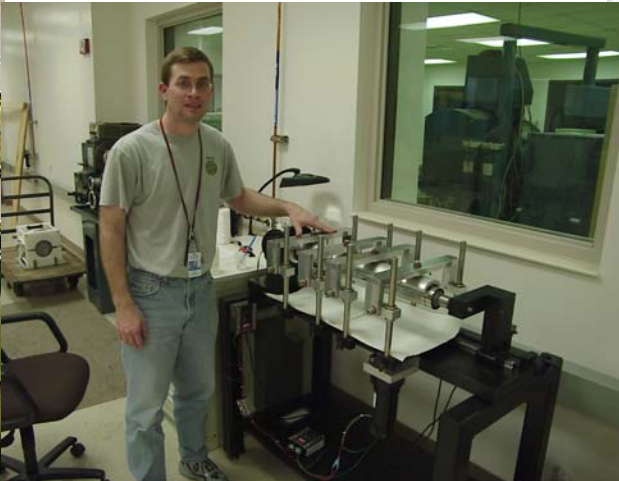
# HERO Program

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- **HERO, for High Energy Replicated Optics is an evolutionary balloon program.**
  - **Utilizes in-house-fabricated hard-x-ray mirrors *plus* supporting x-ray detectors, gondola and pointing system.**
  - **Optic design philosophy :**
    - Utilize a large number of shallow-graze-angle, iridium-coated full-shell mirrors
    - Obtain significant collecting area by using narrow-aspect ratio mirror shells (ie large length to diameter ratio), by nesting many thin shells and by using multiple mirror modules.
  - **Mirrors are made via an electroformed-nickel-replication process. Thin shells are electrodeposited onto a super-polished and figured master (mandrel) from which they are later released by cooling.**
    - Permits multiple high-quality copies to be made from a single master.
    - Provides high-angular-resolution optics
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# Mirror Shell Production

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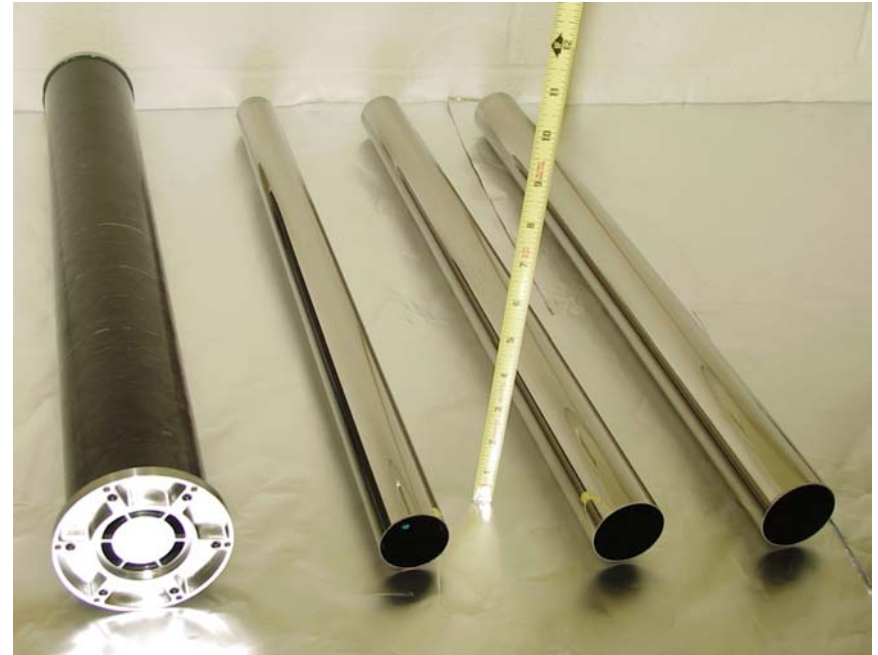


# Current Flight Configuration

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- **Two mirror modules**

- Each module has 3 nested shells
- 40, 44, and 48 mm diameter
- 3 meter focal length
- 2 cm<sup>2</sup> collecting area each
- 45 arcsec module HPD
- Energy cutoff = 50 keV
- FOV (30 keV) = 9 arcmin diam (50%)



- **Mirror Shells**

- 61 cm long
  - 0.25 mm thick glassy nickel alloy
  - Iridium coated
  - Half Power Diameter (HPD) of 30 arcsec
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# Current Flight Configuration

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- **Detectors:** *Imaging Gas Scintillation Proportional Counters*

Sensitive Area: Approximately 20 cm<sup>2</sup>

Fill Gas: 50 mm of Xenon + Helium (96/4) at 10<sup>6</sup> Pa

Entrance Window: 3.2 mm Beryllium

Light Emitting Region: 4.0 mm deep

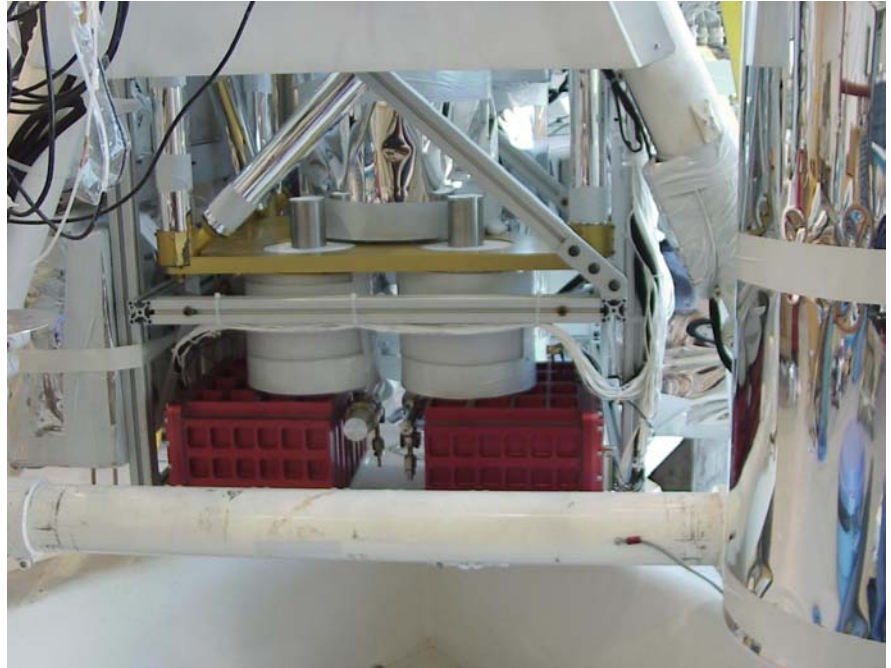
Exit Window: 7.0 mm of Suprasil

Phototube: Hamamatsu 4268, position sensitive,  
quartz window

Quantum Efficiency: 99% @ 40 keV, 89% @ 60 keV

Energy Resolution (FWHM): 5% (30 keV), 4% (60 keV)

Position Resolution (FWHM): 420 $\mu$ m (15-25keV); 330 $\mu$ m (25-35 keV); 400  $\mu$ m (35-45 keV)

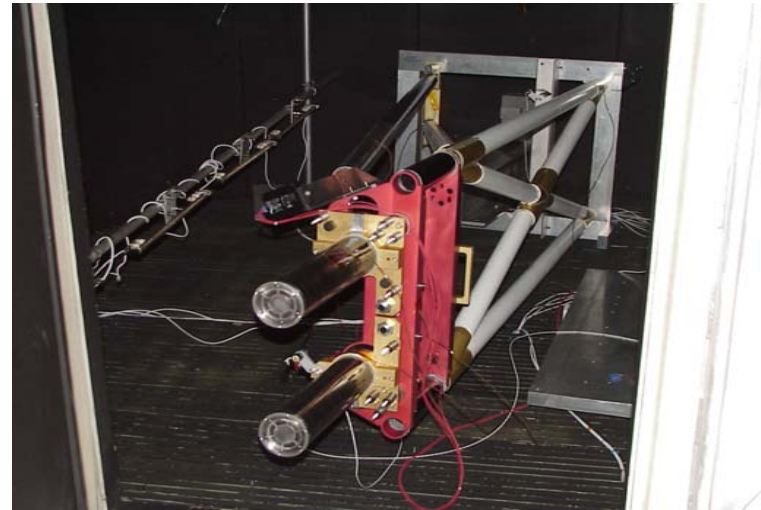


# Current Flight Configuration

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- **Optical Bench**

- Carbon fiber with individual tip/tilt mount for each mirror module
- In-flight laser alignment monitoring system

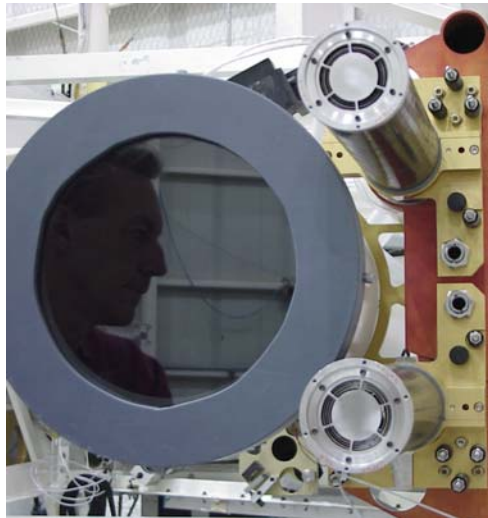


- **Gondola**

- Alt/Azimuth mount for optical bench
  - Coarse pointing system based on global positioning system
  - Fine inertial system based on gyroscopes updated from a day/night star camera
  - Autonomous operation from on-board star catalogs
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# Gondola Assembly

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# The Launch

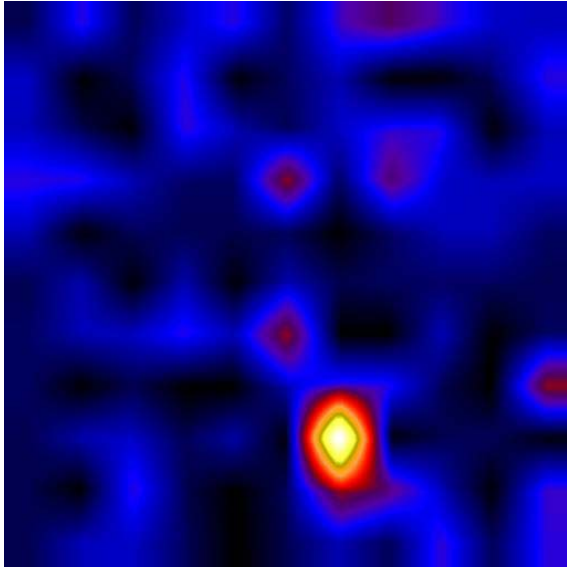
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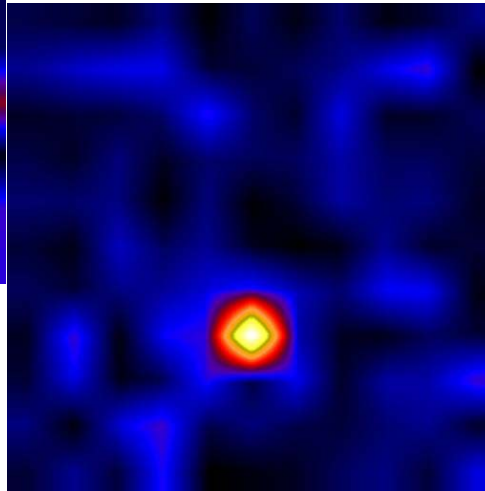


# Images

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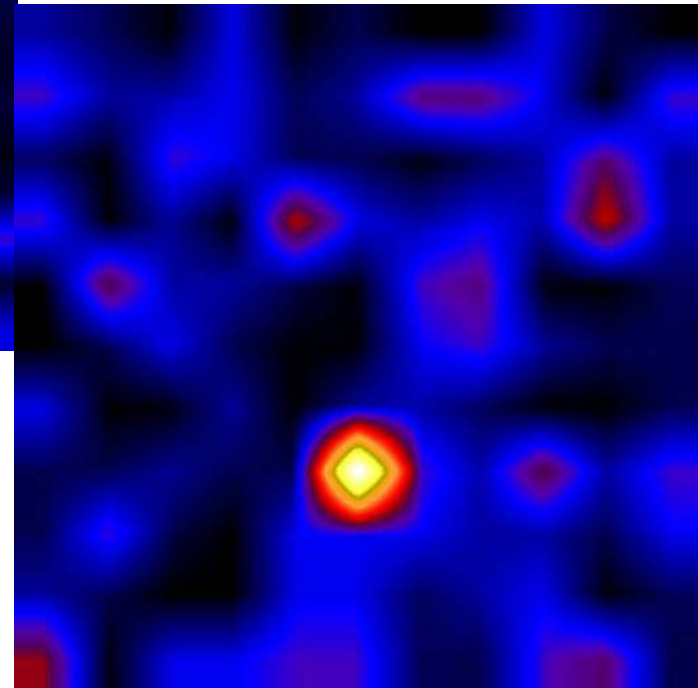


Crab



Cygnus X-1

GRS 1915



# Flight Performance

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- **Payload launched on 23 May, 2001, from Fort Sumner New Mexico**
    - ~ 17 hours at float
    - 39km altitude day, 37 km night
    - All systems worked well
    - Gondola held pointing to < 30 arcsec (33% of time), 60 arcsec (60%)
    - Aspect derived to 8 arcsec day and night for image reconstruction
    - Cyg X-1, Crab and GRS1915 seen at  $8\sigma$ ,  $8\sigma$  and  $6\sigma$  over 20-45 keV
    - ***First hard-x-ray focused images***
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# Future Plans

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- **Modify Gondola for 6-m focal length optics**
    - Longer focal length will give greater collecting area and improved high-energy response
  - **Increase module size**
    - 15 shells ranging from 50 to 95 mm diameter
    - 12 mandrels already in production (first [94 mm] ..metrology gave HPD 10 arcsec core)
    - Effective area 12.5 cm<sup>2</sup> (40 keV) 7.5 cm<sup>2</sup> (60 keV)
  - **Improve angular resolution**
    - 30 arcsec HPD (already demonstrated up to 60 keV)
    - 10 – 15 arcsec HPD goal
  - **Increase number of modules**
    - 8 by 2003
    - 16 by 2004 – 200 cm<sup>2</sup> at 40 keV. ...this will give a sensitivity 1000 times greater HEAO-3 survey
  - **Investigate adding extra outer shells with multilayer coatings**
    - Extend energy range beyond Iridium K edge (76 keV)
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# New High-Resolution 6-m-Focal Length Mirror Shells

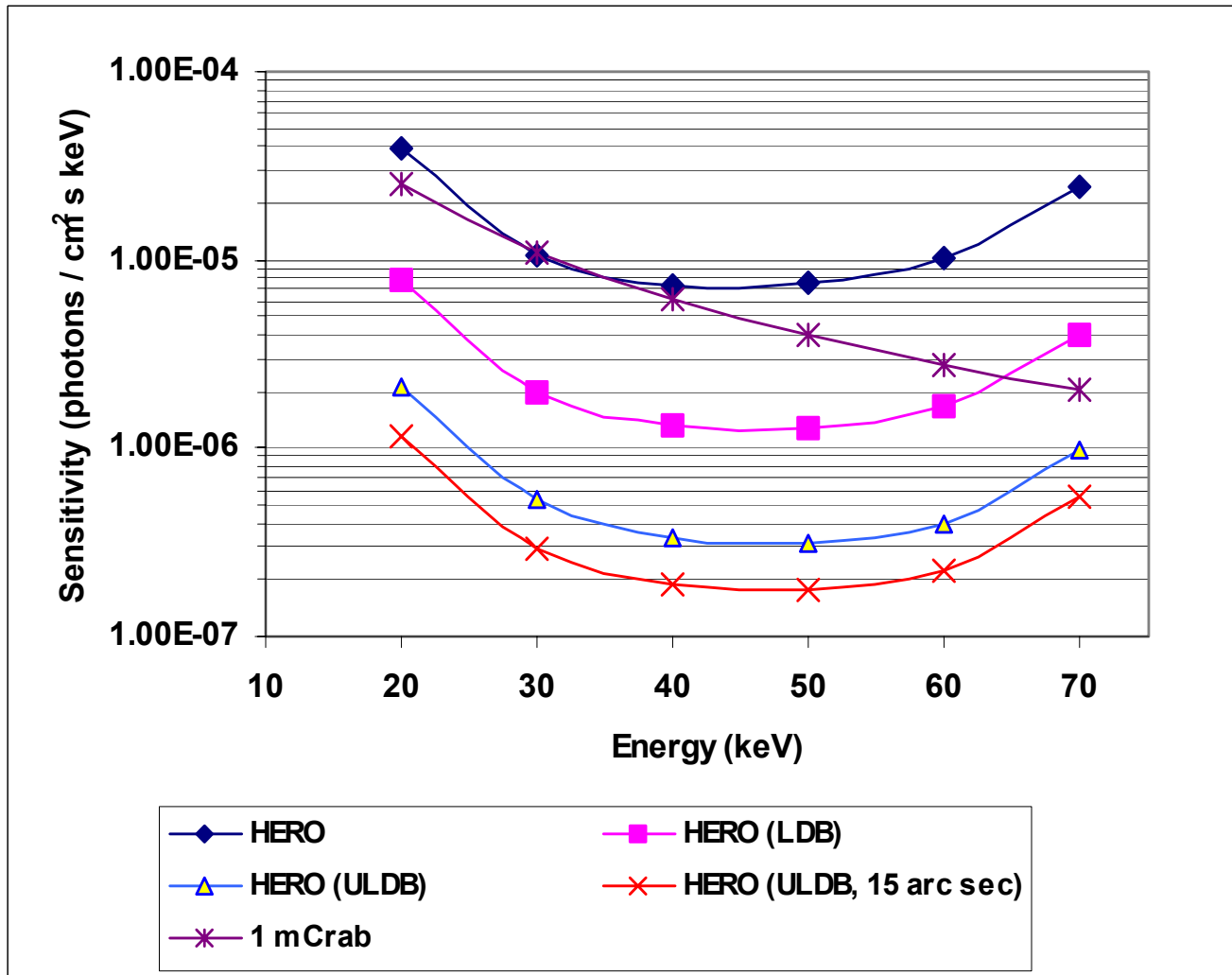
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- New mirror shells from 10 arcsec mandrels



- Above shows new outer HERO shell (94mm) alongside spring flight module.
  - Uncoated shell was recently measured to have 16 arcsec HPD (20 keV) with simple support.
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# 5- $\sigma$ Sensitivities (16 Mirror Modules)



# Application to Con-X HXT

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- Some developments of relevance to HXT :
    - Low cost mandrel production (<10k / mandrel)
    - High strength alloys permit very thin shells
    - Release under water/alcohol provides low-stress separation of thin shells
      - *High angular resolution demonstrated*
  - To meet the (original) weight budget of HXT requires shells ranging from 12 cm to 40 cm diameter, and thickness from 0.1mm to 0.15 mm.
    - HERO shells are 0.25 mm thick, 9.4 cm diameter (outer)
      - *But 50 cm diameter, 0.15-mm-thick shells have also been fabricated*
  - MSFC will provide replicated shells to SAO for multilayer coatings. These will then be tested at MSFC .
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# Large-Mirror-Shell Development

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- Figure at right shows a 0.15 mm thick, 50 cm diameter shell on the metrology station.
- Metrology of figure predicted a geometric core of 38 arcsec HPD
- After stress relief this improved to 19 arcsec HPD.

